

The LIFE initiative – atmospheric characterization of terrestrial exoplanets in the mid-infrared with a large space-based nulling interferometer

Authors:

Sascha P. Quanz
ETH Zurich

ETH zürich



CENTRE FOR
ORIGIN &
PREVALENCE OF
LIFE

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National Centre of Competence in Research

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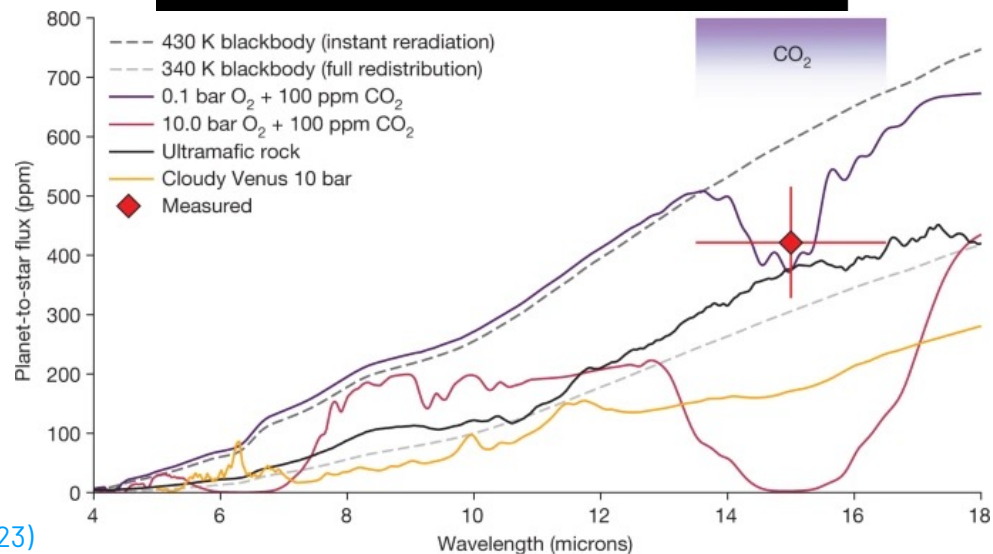
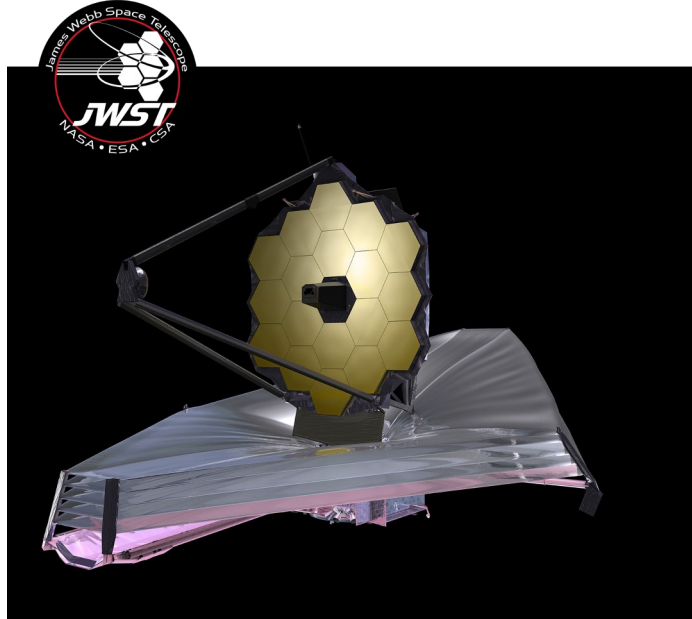
SF2A workshop – S12
20/21 June, 2023
Strasbourg

The LIFE initiative seeks to develop the scientific context, the technology, and a roadmap for an ambitious mid-infrared space mission that investigates the atmospheric properties of a large sample of terrestrial exoplanets.

LIFE's primary mission objective:

Determining the occurrence rate of life – as we now it – on nearby exoplanets

First steps are taken by JWST and Ariel

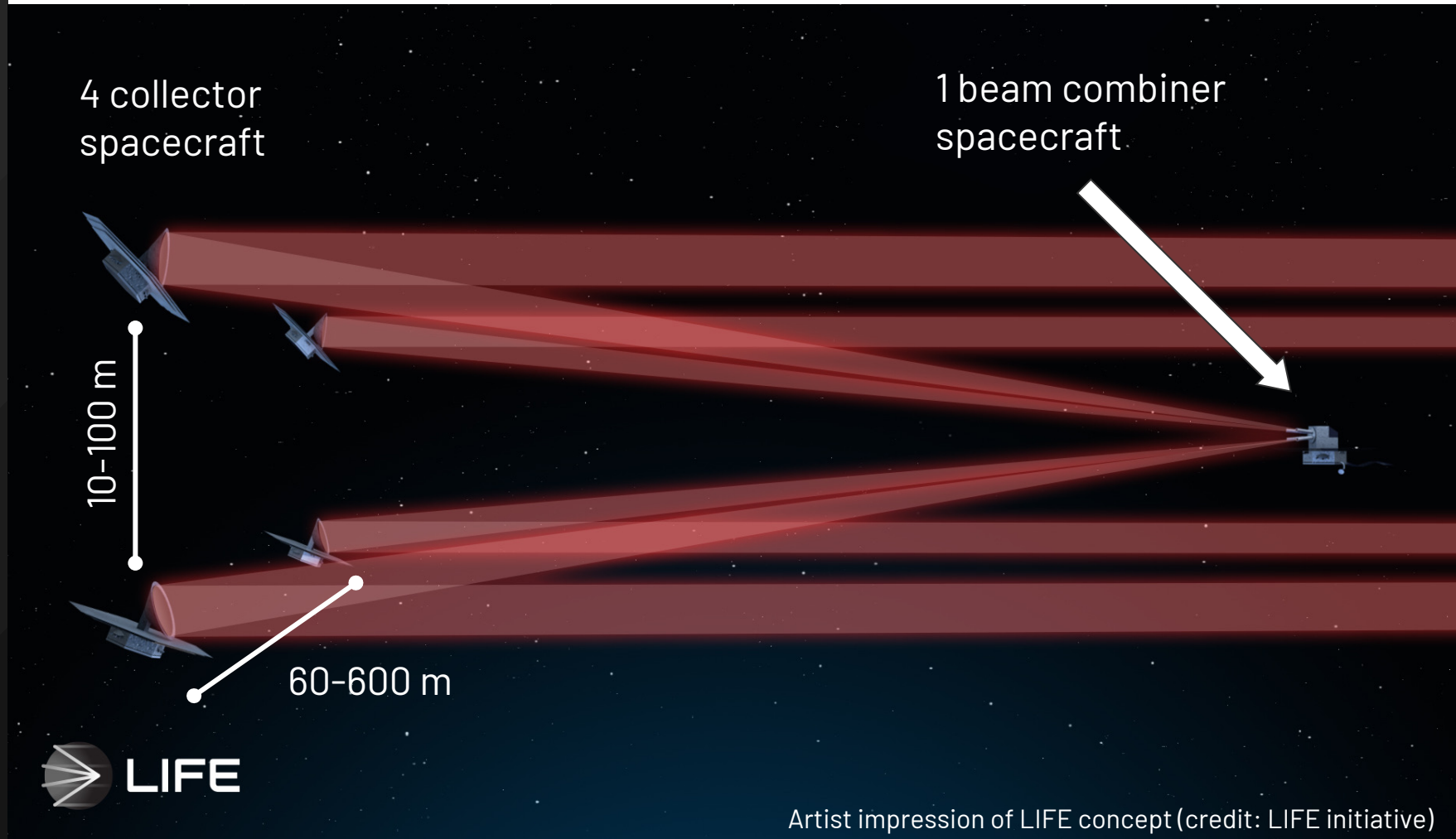


“A long term scientific objective is to characterize the whole range of exoplanets, including, of course, potentially habitable ones. ARIEL would act as a pathfinder **for future, even more ambitious campaigns.**”

ARIEL Assessment Study Report (Yellow Book)

The LIFE mission

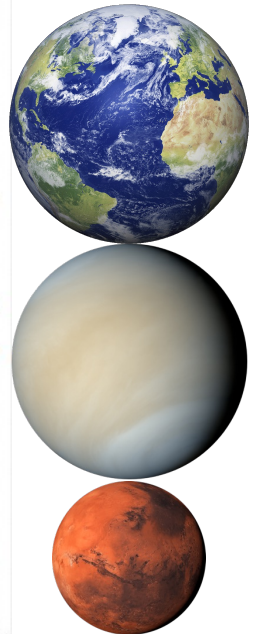
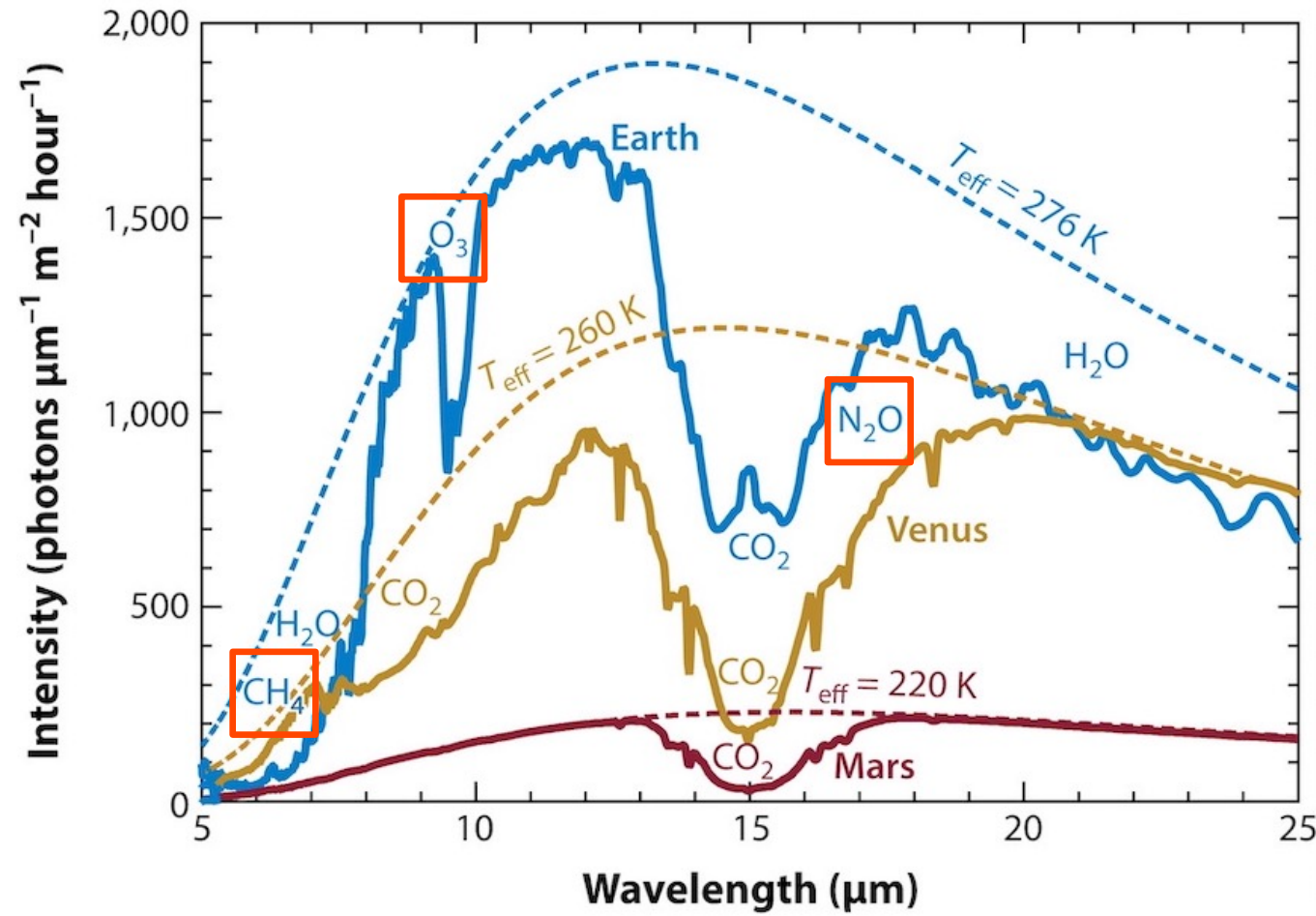
- ...is a space-based formation-flying mid-infrared (nulling) interferometer in L2 with a nominal mission lifetime of at least 5 years
- ...consists of 4 collector spacecraft separated by tens to hundreds of meters and a beam combiner spacecraft
- ...will survey hundreds of nearby stars within 25 parsec, discover hundreds of new exoplanets, and investigate the atmospheres of dozens of them



Investigating other worlds

- LIFE's wavelength range is chosen to cover the peak of the thermal emission of temperate terrestrial planets
- This wavelength range features absorption bands of major atmospheric constituents including biosignatures such as ozone (O_3), methane (CH_4) and nitrous oxide (N_2O)

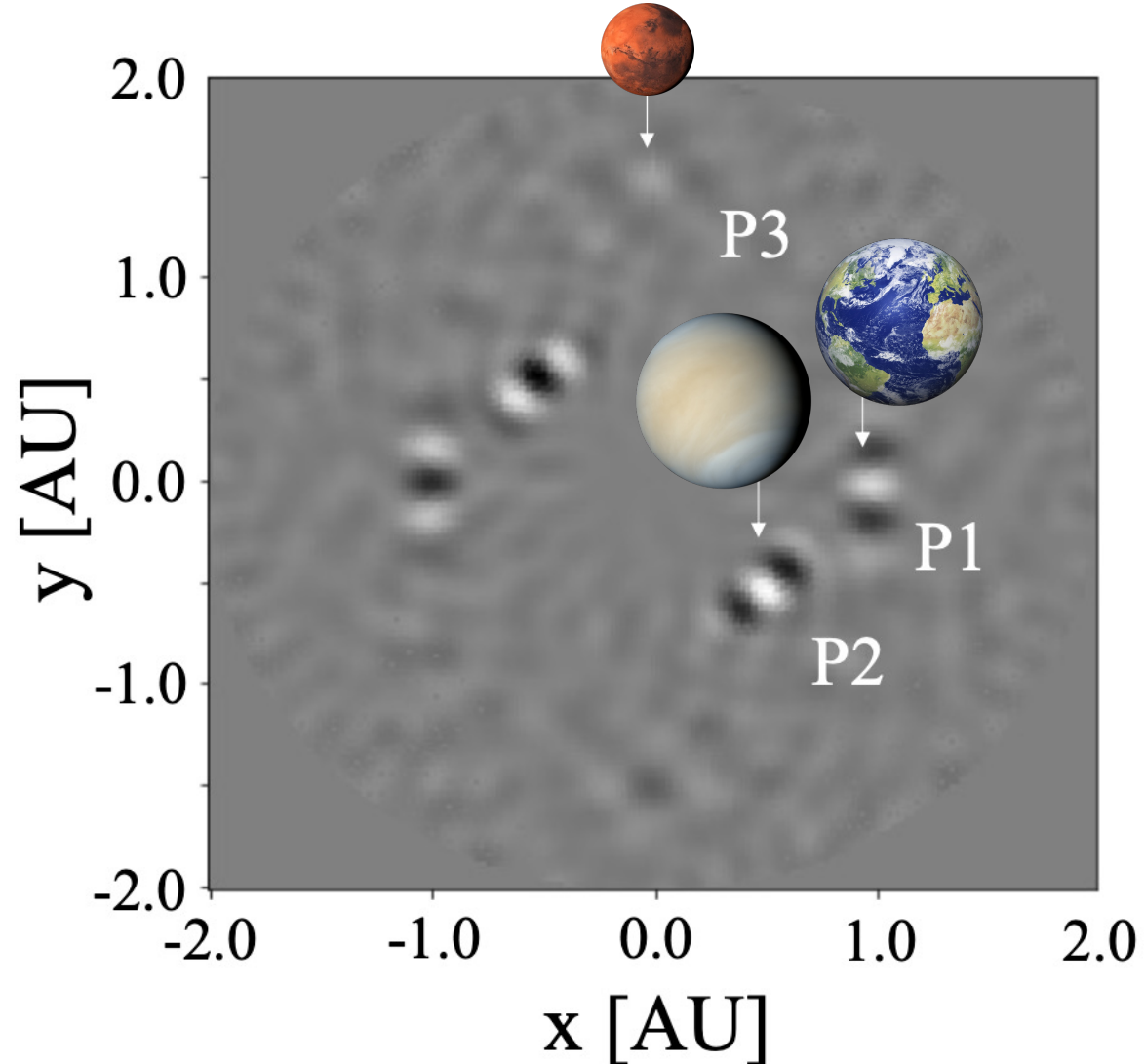
Emission spectra of terrestrial planets in our Solar System



Investigating other worlds

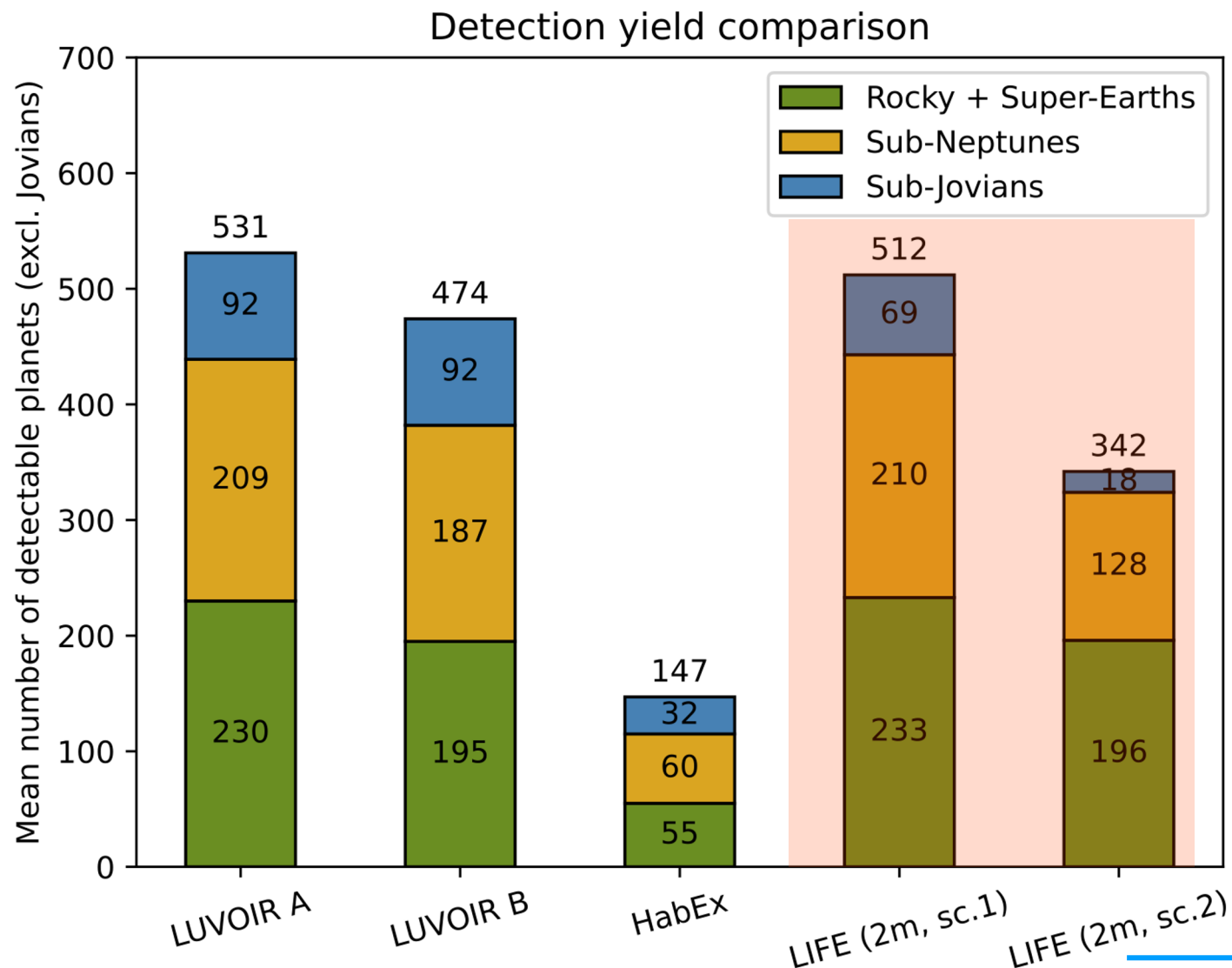
- LIFE's wavelength range is chosen to cover the peak of the thermal emission of temperate terrestrial planets
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Reconstructed LIFE image of Solar System analog at 10 pc



LIFE: Exoplanet Detection Yield Estimates

Assuming an initial search phase of the mission of 2.5 years



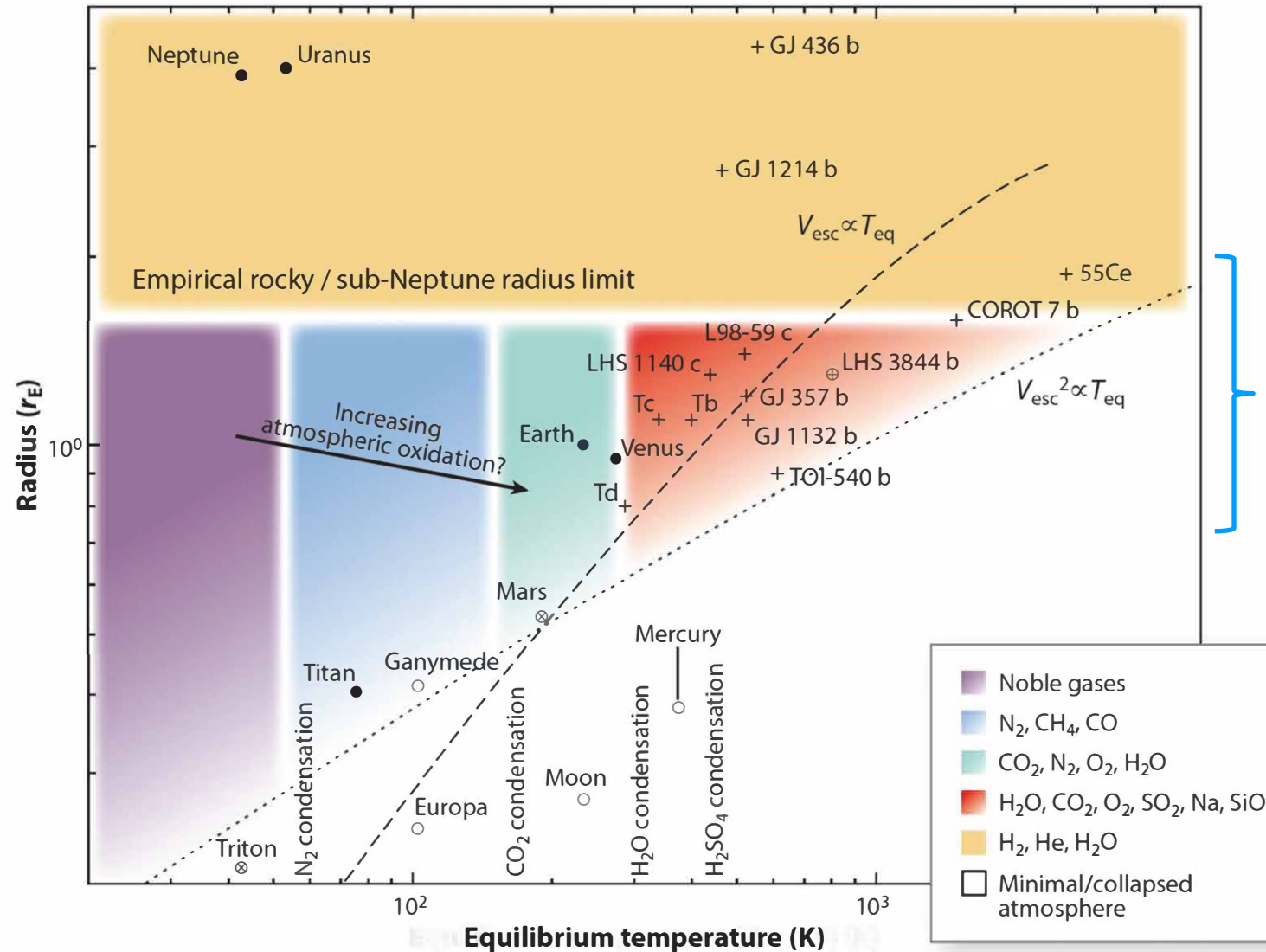
- Expected detection yields are similar to large future NASA flagship concepts
- Monte Carlo simulations based on Kepler statistics (SAG13) and stars within ~20 pc

2 scenarios:

Maximizing **total number** of planets vs.
maximizing **rocky planets in habitable zone**

LIFE provides unprecedented discovery potential

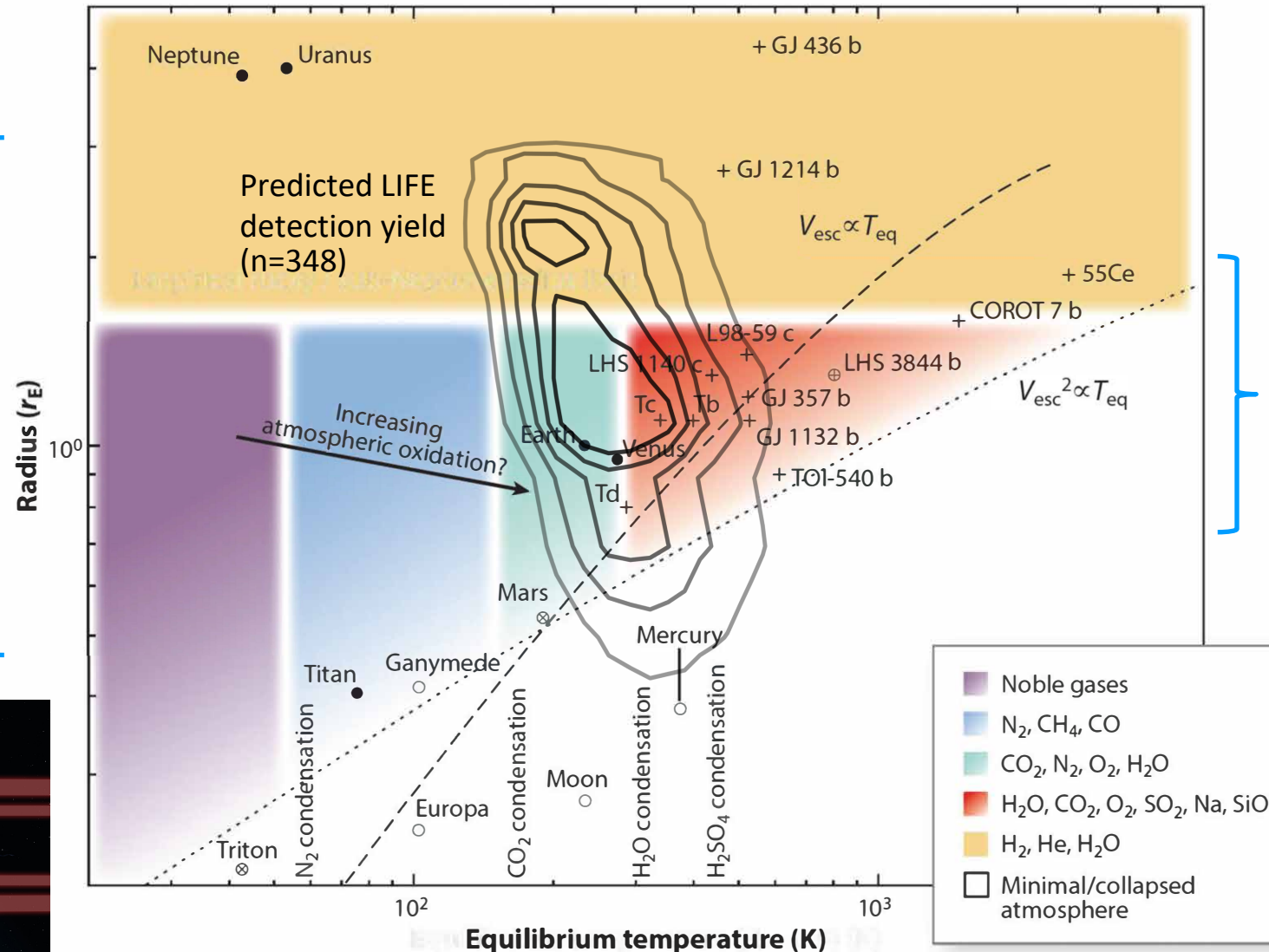
LIFE discovery space vs. rocky exoplanets with JWST



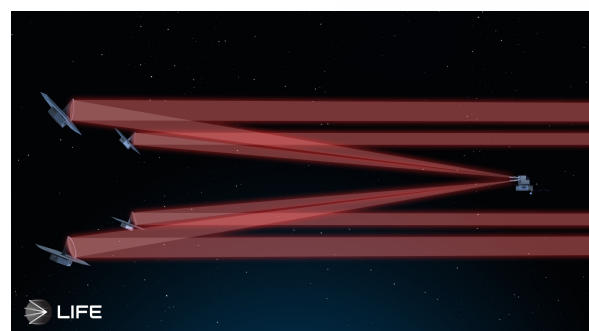
LIFE provides unprecedented discovery potential

LIFE discovery space vs. rocky exoplanets with JWST

Primary discovery space of LIFE mission

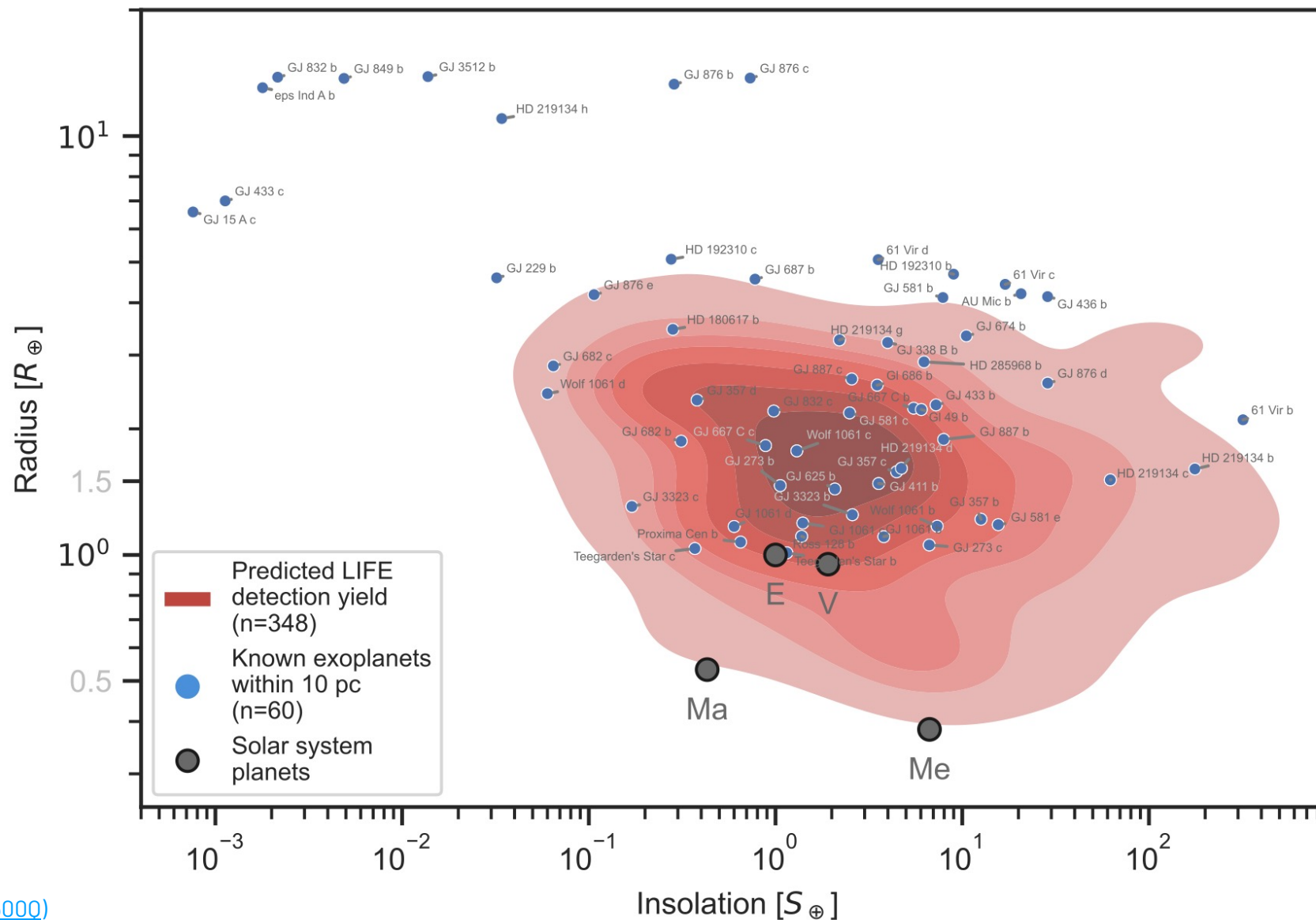


Crosses: All known "rocky-ish" exoplanets to be observed with JWST



LIFE is a characterization mission from day 1

LIFE discovery space vs. known exoplanets within 10 pc



Exoplanet characterization: the mid-infrared advantage

In contrast to a reflected light mission, LIFE will...



...directly constrain the **pressure-temperature structure** of exoplanet atmospheres



...access (multiple) atmospheric absorption bands of **major molecules** such as H_2O , CO_2 , and CO as well as collision induced absorption from N_2 and O_2



...search for numerous **atmospheric biosignatures** in the context of terrestrial exoplanets and gas dominated Super-Earths (e.g., O_3 and CH_4 , but also N_2O , PH_3 , NH_3 , and C_5H_8)



...constrain directly **the effective temperature** of exoplanets and provide access to their **radii**



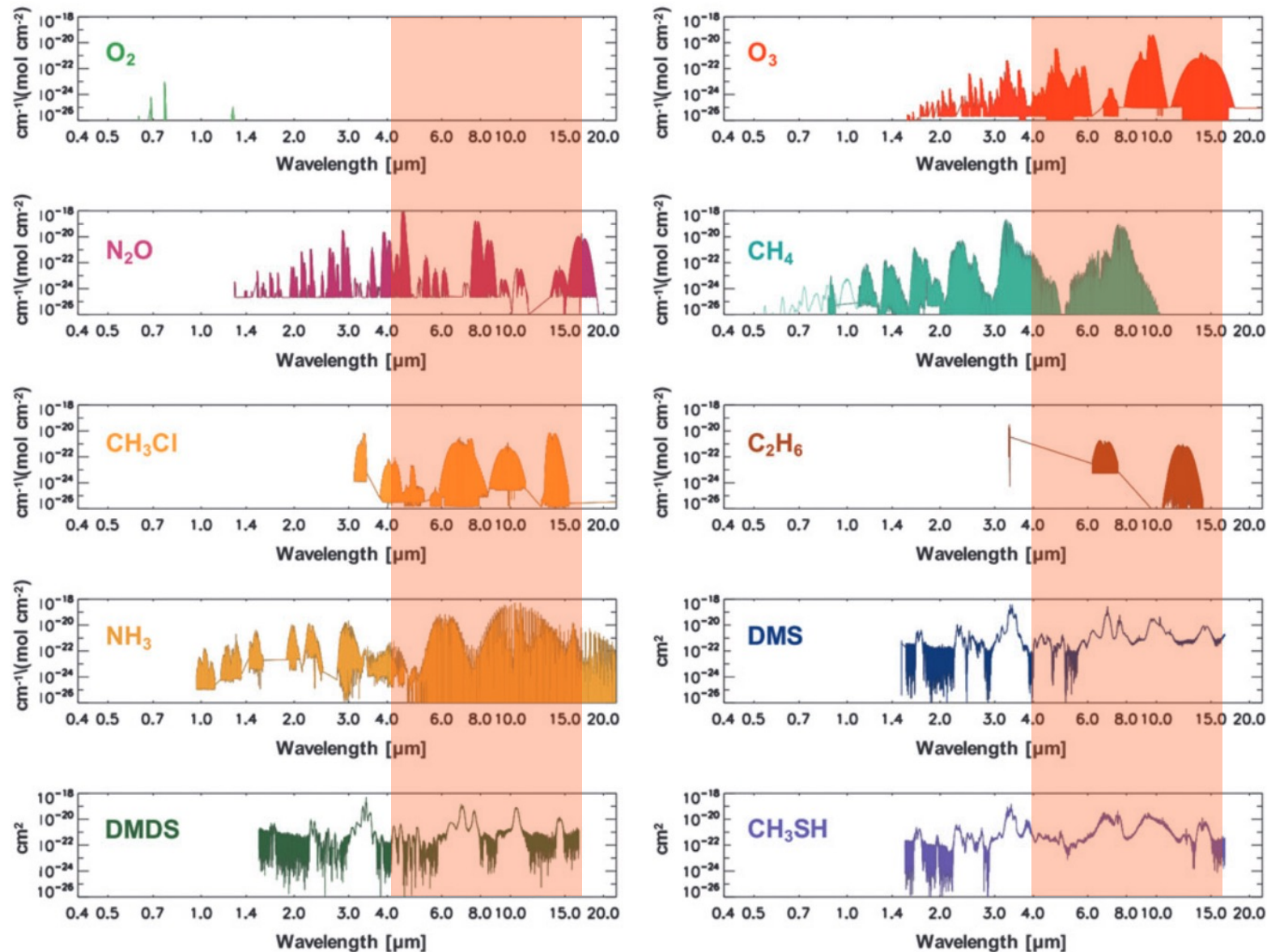
...deliver a higher detection yield during search phase as it is **less affected by the orbital phase function** of the exoplanets' emission compared to reflected light missions



...immediately **start observing already known small, temperate exoplanets** around nearby M-stars

Biosignature detection: the mid-infrared advantage

Many atmospheric biosignatures have absorption bands in the LIFE wavelength range



 LIFE wavelength range









LIFE-related paper series is a growing success

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Astrophysics

Large Interferometer For Exoplanets (LIFE)

I. Improved exoplanet detection yield estimates for a large mid-infrared space-interferometer mission










S. P. Quanz^{1,2} , M. Ottiger¹, E. Fontanet¹, J. Kammerer^{3,4,22} , F. Menti¹, F. Dannert¹ , A. Gheorghe¹, O. Absil⁵ , V. S. Airapetian⁶, E. Alei^{1,2} , R. Allart⁷ , D. Angerhausen^{1,2}, S. Blumenthal⁸, L. A. Buchhave⁹, J. Cabrera¹⁰, Ó. Carrión-González¹¹ , G. Chauvin¹² , W. C. Danchi⁶, C. Dandumont¹³, D. Defrère¹⁴, C. Dorn¹⁵, D. Ehrenreich¹⁶, S. Ertel^{17,18}, M. Fridlund^{19,20}, A. García Muñoz¹¹, C. Gascón²¹, J. H. Girard²², A. Glauser¹, J. L. Grenfell¹⁰, G. Guidi^{1,2}, J. Hagelberg¹⁶, R. Helled¹⁵, M. J. Ireland⁴, M. Janson²³, R. K. Kopparapu⁶, J. Korth²⁴, T. Kozakis⁹, S. Kraus²⁵, A. Léger²⁶, L. Leedjävry²⁷, T. Lichtenberg⁸, J. Lillo-Box²⁸, H. Linz²⁹, R. Liseau²⁰, J. Loicq¹³, V. Mahendra³⁰, F. Malbet¹⁰, J. Mathew³, B. Mennesson³¹, M. R. Meyer³², L. Mishra^{33,16,2}, K. Molaverdikhani^{29,34}, L. Noack³⁵, A. V. Oza^{31,33}, E. Pallé^{26,37}, H. Parviainen^{36,37}, A. Quirrenbach³⁸, H. Rauer¹⁰, I. Ribas^{21,38}, M. Rice³⁹, A. Romagnolo⁴⁰, S. Rugheimer⁸, E. W. Schwieterman⁴¹, E. Serabyn³¹, S. Sharma⁴², K. G. Stassun⁴³, J. Szulágyi¹, H. S. Wang^{1,2}, F. Wunderlich¹⁰, M. C. Wyatt⁴⁴, and the LIFE Collaboration⁴⁵

A&A 664, A22 (2022)
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Large Interferometer For Exoplanets (LIFE)

II. Signal simulation, signal extraction, and fundamental exoplanet parameters from single-epoch observations










Felix A. Dannert^{1,2*} , Maurice Ottiger^{1,*,}, Sascha P. Quanz^{1,2} , Romain Laugier³, Emile Fontanet¹ , Adrian Gheorghe¹, Olivier Absil^{4,***} , Colin Dandumont⁵, Denis Defrère³ , Carlos Gascón⁶, Adrian M. Glauser¹ , Jens Kammerer⁷ , Tim Lichtenberg⁸ , Hendrik Linz⁹ , Jérôme Loicq^{5,10}, and the LIFE collaboration***

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Large Interferometer For Exoplanets (LIFE)

III. Spectral resolution, wavelength range, and sensitivity requirements based on atmospheric retrieval analyses of an exo-Earth

B. S. Konrad^{1,2} , E. Alei^{1,2} , S. P. Quanz^{1,2} , D. Angerhausen^{1,2,3} , Ó. Carrión-González⁴ , J. J. Fortney⁵ , J. L. Grenfell⁶, D. Kitzmann⁷ , P. Mollière⁸ , S. Rugheimer⁹ , F. Wunderlich⁶, and the LIFE Collaboration*

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Large Interferometer For Exoplanets (LIFE)

IV. Ideal kernel-nulling array architectures for a space-based mid-infrared nulling interferometer






Jonah T. Hansen , Michael J. Ireland, and the LIFE Collaboration*

A&A 665, A106 (2022)
<https://doi.org/10.1051/0004-6361/202243760>
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Large Interferometer For Exoplanets (LIFE)

V. Diagnostic potential of a mid-infrared space interferometer for studying Earth analogs


Eleonora Alei^{1,2} , Björn S. Konrad^{1,2} , Daniel Angerhausen^{1,2,3} , John Lee Grenfell⁴, Paul Mollière⁵, Sascha P. Quanz^{1,2} , Sarah Rugheimer⁶ , Fabian Wunderlich⁴, and the LIFE Collaboration*

A&A 668, A52 (2022)
<https://doi.org/10.1051/0004-6361/202243846>
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Large Interferometer For Exoplanets (LIFE)

VI. Detecting rocky exoplanets in the habitable zones of Sun-like stars

Jens Kammerer¹ , Sascha P. Quanz^{2,3} , Felix Dannert² , and the LIFE Collaboration⁴

A&A 670, A57 (2023)
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Large Interferometer For Exoplanets (LIFE)

VII. Practical implementation of a five-telescope kernel-nulling beam combiner with a discussion on instrumental uncertainties and redundancy benefits

Jonah T. Hansen¹ , Michael J. Ireland¹ , Romain Laugier², and the LIFE Collaboration*

A&A 671, A114 (2023)
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Imaging of exocomets with infrared interferometry

Markus Janson¹, Jayshil Patel¹, Simon C. Ringqvist¹, Cicero Lu², Isabel Rebollido³, Tim Lichtenberg^{4,5}, Alexis Brandeker¹, Daniel Angerhausen^{6,7,8}, and Lena Noack⁹

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Large Interferometer For Exoplanets (LIFE)

IX. Assessing the impact of clouds on atmospheric retrievals at mid-infrared wavelengths with a Venus-twin exoplanet*

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DOI: 10.1089/ast.2022.0010

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Large Interferometer for Exoplanets: VIII. Where Is the Phosphine? Observing Exoplanetary PH₃ with a Space-Based Mid-Infrared Nulling Interferometer

Daniel Angerhausen,^{1–3} Maurice Ottiger,¹ Felix Dannert,¹ Yamila Miguel,^{4,5} Clara Sousa-Silva,^{6,7} Jens Kammerer,⁸ Franziska Menti,¹ Eleonora Alei,^{1,2} Björn S. Konrad,^{1,2} Haiyang S. Wang,^{1,2} Sascha P. Quanz,^{1,2}, and The LIFE Collaboration⁹

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The Nulling Interferometer Cryogenic Experiment: I

Mohanakrishna Ranganathan, Adrian Glauser, Thomas Birbacher, Adrian Gheorghe, Sascha Quanz

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A sub-nanometer long-term stable heterodyne laser metrology system for the Nulling Interferometry Cryogenic Experiment

Thomas Birbacher, Adrian Glauser, Mohanakrishna Ranganathan, Sascha Quanz

Mohanakrishna Ranganathan, Adrian M. Glauser, Thomas Birbacher, Adrian A. Gheorghe, Sascha P. Quanz, "The Nulling Interferometry Cryogenic Experiment: I," Proc. SPIE 12183, Optical and Infrared Interferometry and Imaging VIII, 121830L (26 August 2022); doi: 10.1117/12.2629514
Event: SPIE Astronomical Telescopes + Instrumentation, 2022, Montréal, Québec, Canada

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Thomas Birbacher, Adrian M. Glauser, Mohanakrishna Ranganathan, Sascha P. Quanz, "A sub-nanometer long-term stable heterodyne laser metrology system for the Nulling Interferometry Cryogenic Experiment," Proc. SPIE 12183, Optical and Infrared Interferometry and Imaging VIII, 121831A (29 August 2022); doi: 10.1117/12.2631654
Event: SPIE Astronomical Telescopes + Instrumentation, 2022, Montréal, Québec, Canada

SPIE

Recent progress and ongoing efforts increase technological readiness

Major technological challenges for mid-infrared space-based interferometry are being tackled by various groups

1

Cryogenic nulling

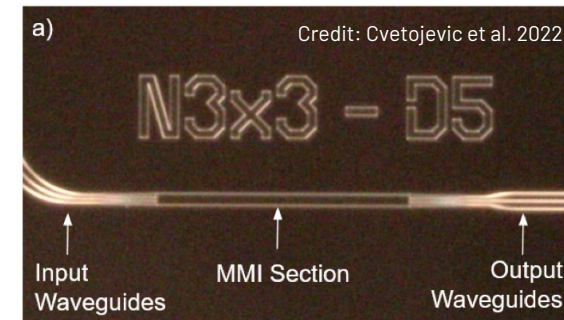
New mid-infrared testbench under construction at ETH Zurich to demonstrate interferometric nulling under realistic conditions



2

Photonics

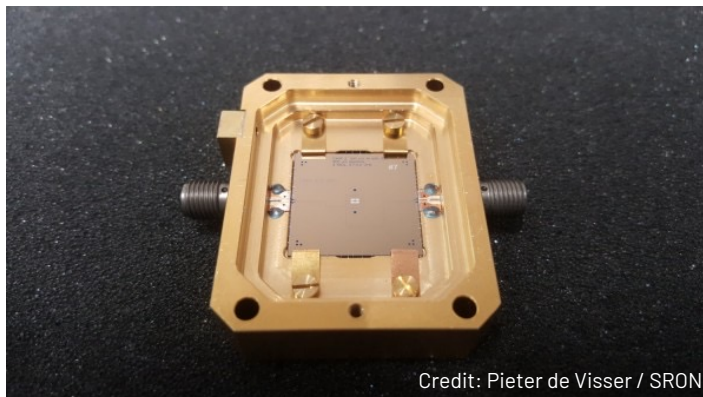
Major breakthroughs in astro-photonics for interferometric nulling at near-infrared wavelengths motivate mid-infrared applications as next step



3

Low-noise detectors

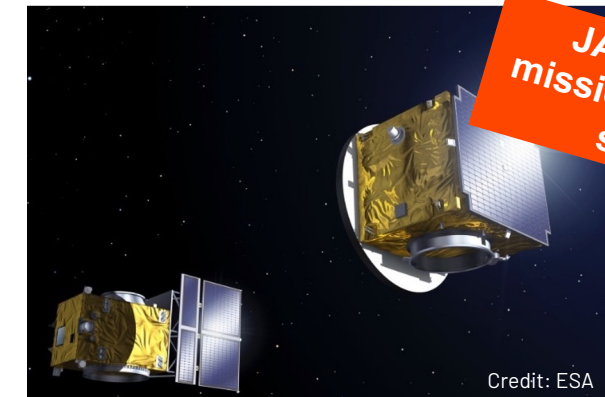
KIDs (Kinetic Inductance Detectors) show excellent performance at sub-mm and near-infrared wavelengths and close-in on mid-infrared regime



4

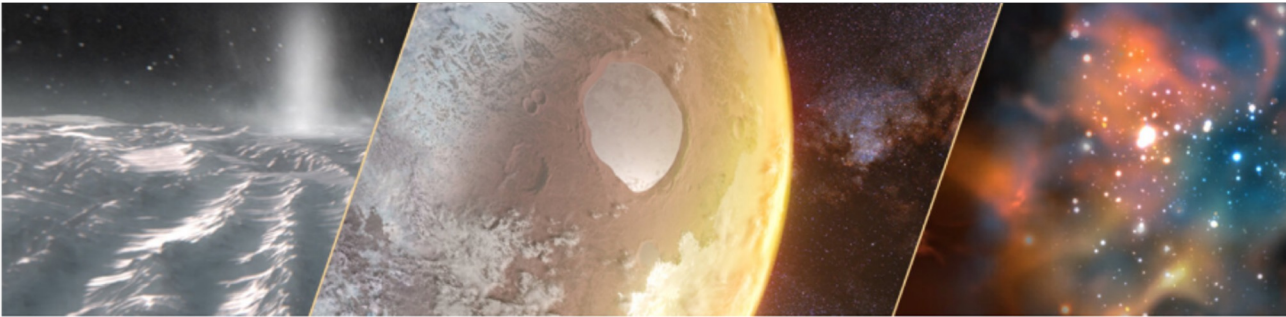
Autonomous formation flying

Various space missions aim to demonstrate high-precision formation flying performance in the coming years including ESA's Proba 3



LIFE: a candidate theme for a future ESA L-class missions

ESA Voyage 2050 – European roadmap for future space exploration



SCIENCE & EXPLORATION

Voyage 2050 sets sail: ESA chooses future science mission themes

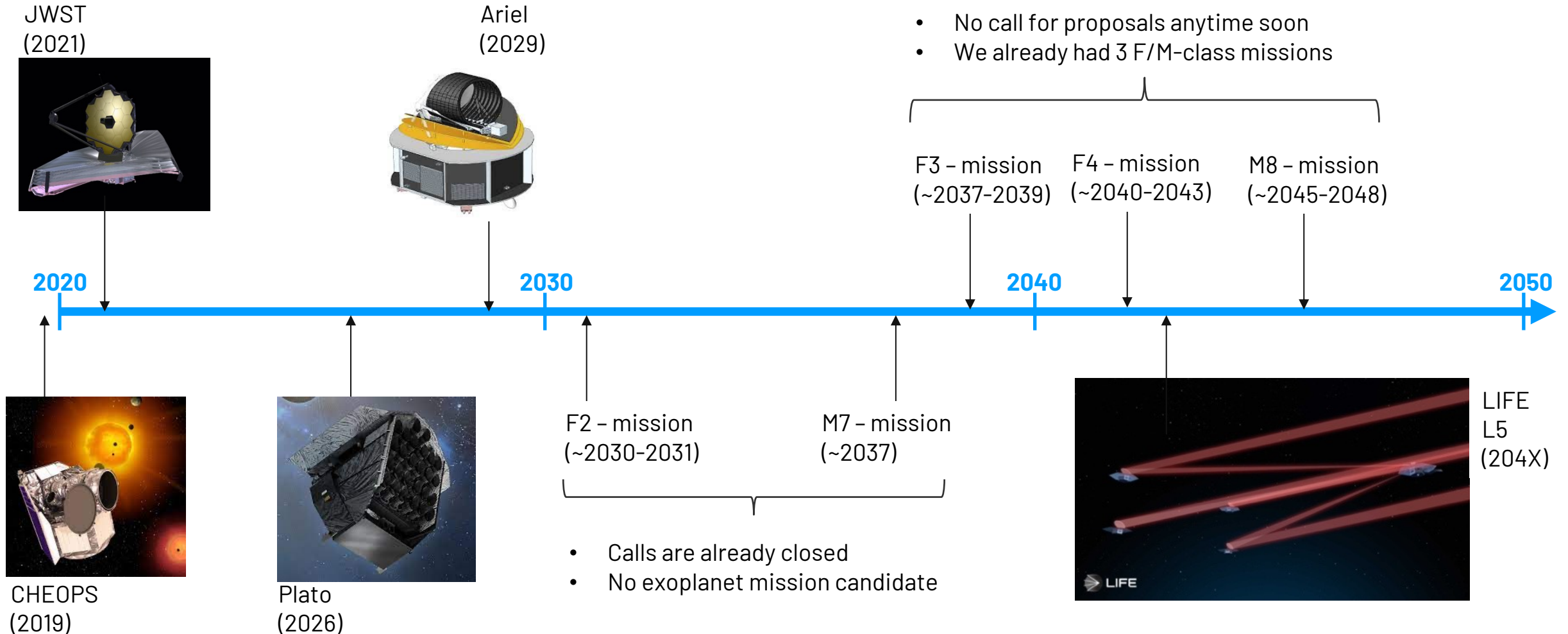
“Therefore, launching a Large mission enabling the characterisation of the **atmosphere of temperate exoplanets in the mid-infrared** should be a top priority for ESA within the Voyage 2050 timeframe.”

“This would give ESA and the European community the opportunity to **solidify its leadership** in the field of exoplanets, [...]”

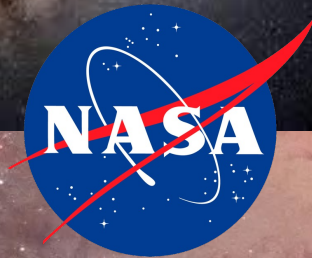
“Being the first to measure a spectrum of the direct thermal emission of a temperate exoplanet in the mid infrared **would be an outstanding breakthrough** that could lead to yet again another paradigm-shifting discovery.”

[ESA Senior Committee Report](#); June 2021

LIFE: a unique opportunity for Europe to lead the way...



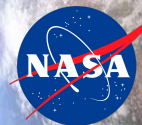
...for the next Copernican revolution with established partners



Lead



Potential
partners

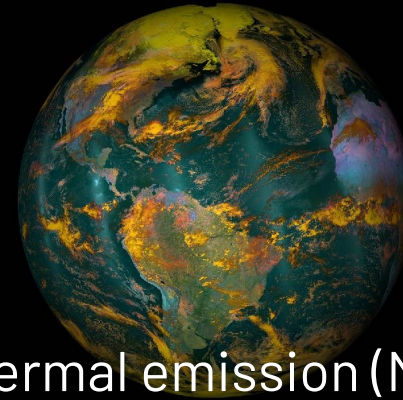


Coordinated efforts will provide a holistic view of Earth-analogs

Synergies between different missions and ground-based telescopes



Reflected light (UV - NIR)



Thermal emission (MIR)

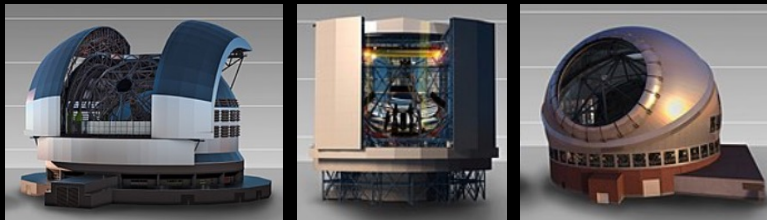
Solar-type
stars

NASA's
HWO



M stars

ELTs

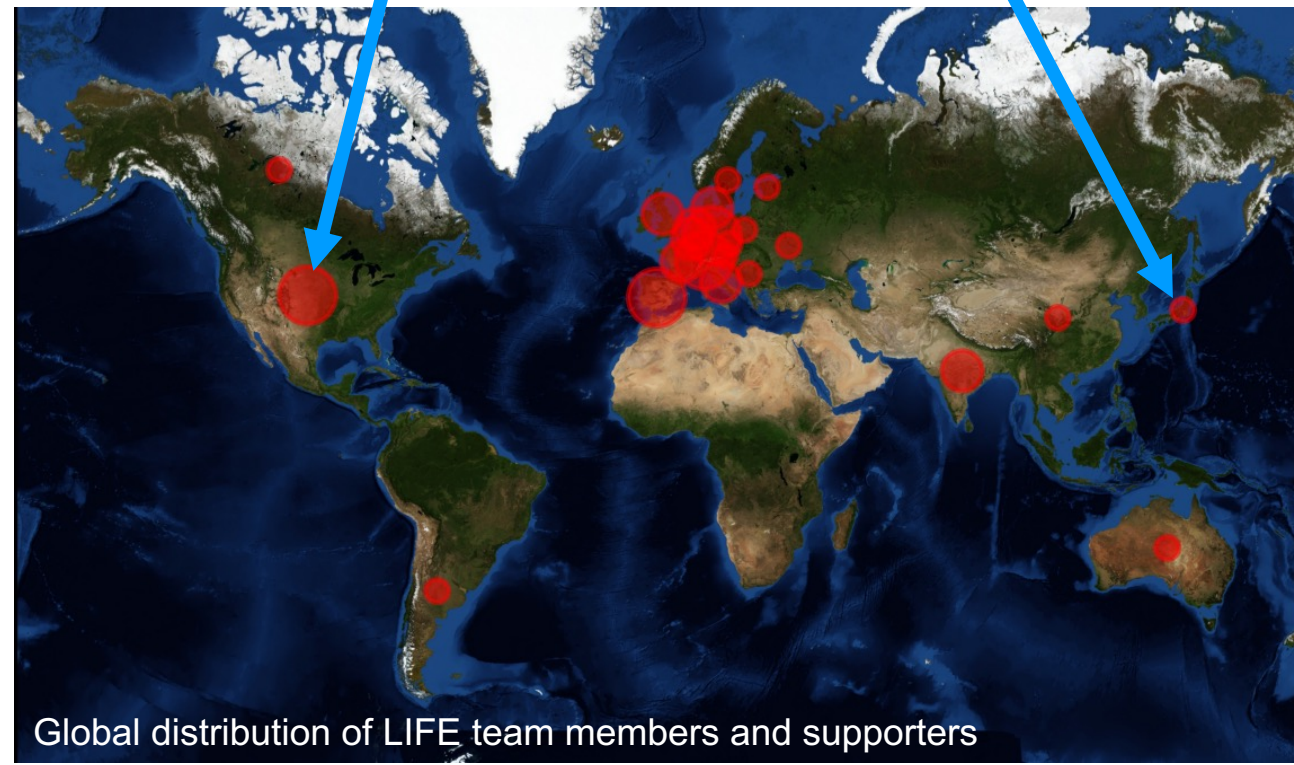


Join the global LIFE initiative!

- LIFE is a European-led but global initiative for a large future exoplanet mission!
- The science theme has been recognized as a potential candidate for an L-class mission within ESA's Science Programme
- R&D for critical components / sub-systems is starting to ramp up, **and we are working towards a mission concept study with our academic and industry partners (kick-off 2024)**
- LIFE is not a closed-club; collaborations / contributions / partnerships at various levels are more than welcome!
- More information:
 - www.life-space-mission.com
 - Sign up for newsletter: life@phys.ethz.ch
 - Follow us: @LIFE_telescope

Collaborations with
NASA/JPL/Goddard on
science and synergies with
future missions

Collaborations with
Japan/JAXA in context of
their SILVIA mission and
other relevant technologies





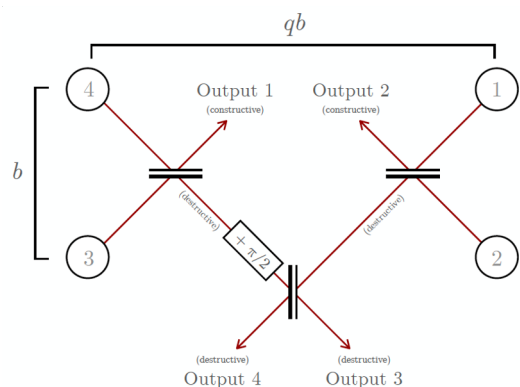
Thank you



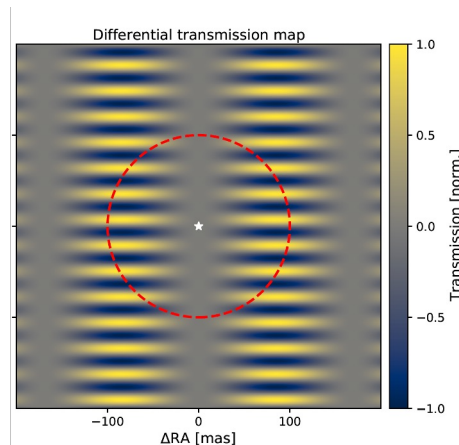
From Source to Detection

Summary Slide

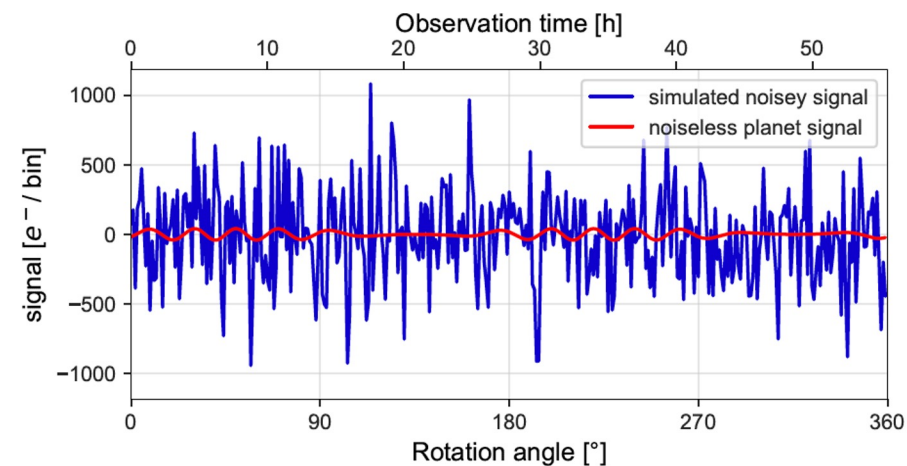
Instrument Model



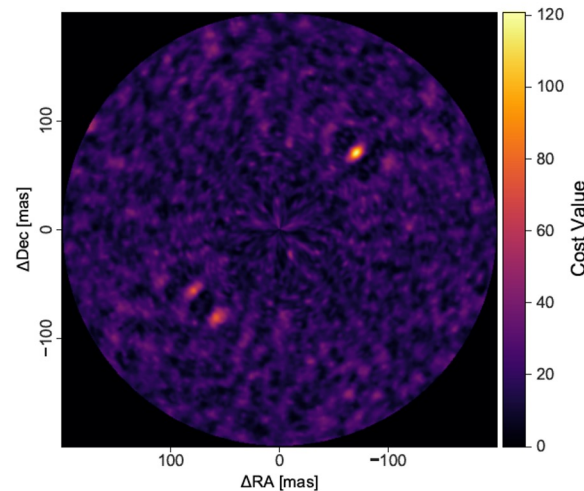
Transmission Map



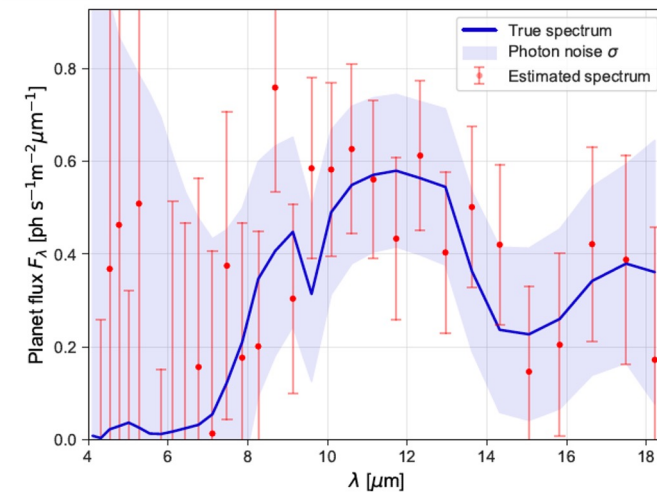
Noisy Time Series



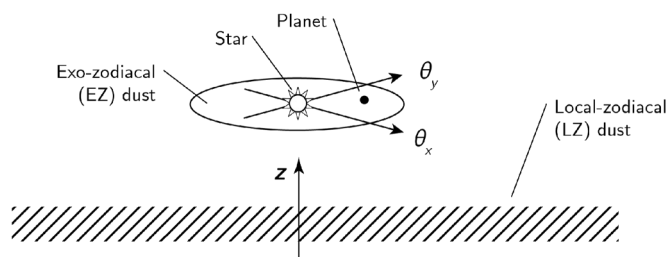
Detection Map



Spectrum



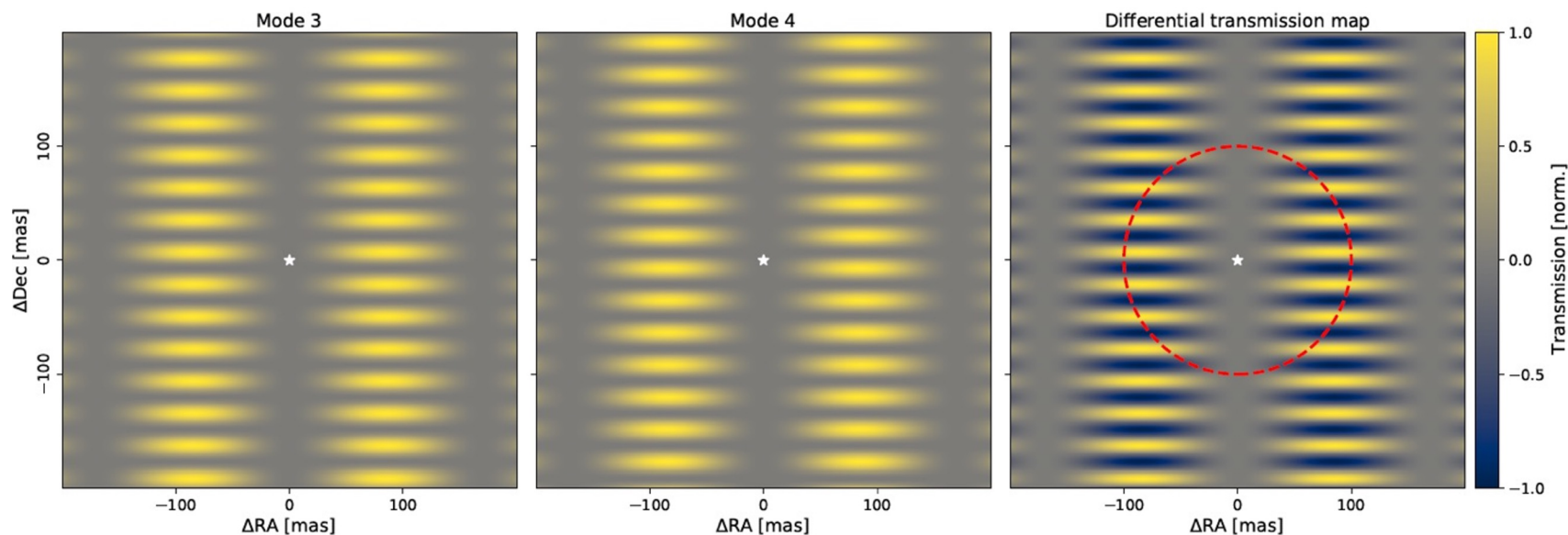
Source Simulation



Lay 2004

Signal Simulation

- Double Bracewell nulling interferometer; there are 2 constructive and 2 destructive outputs
- In one branch, a $\pi/2$ phase shift is introduced to enable the difference map
- Phase Chopping between Outputs 3 & 4 makes instrument less susceptible to perturbation
- Assume fundamental noise limit (ideal instrument)

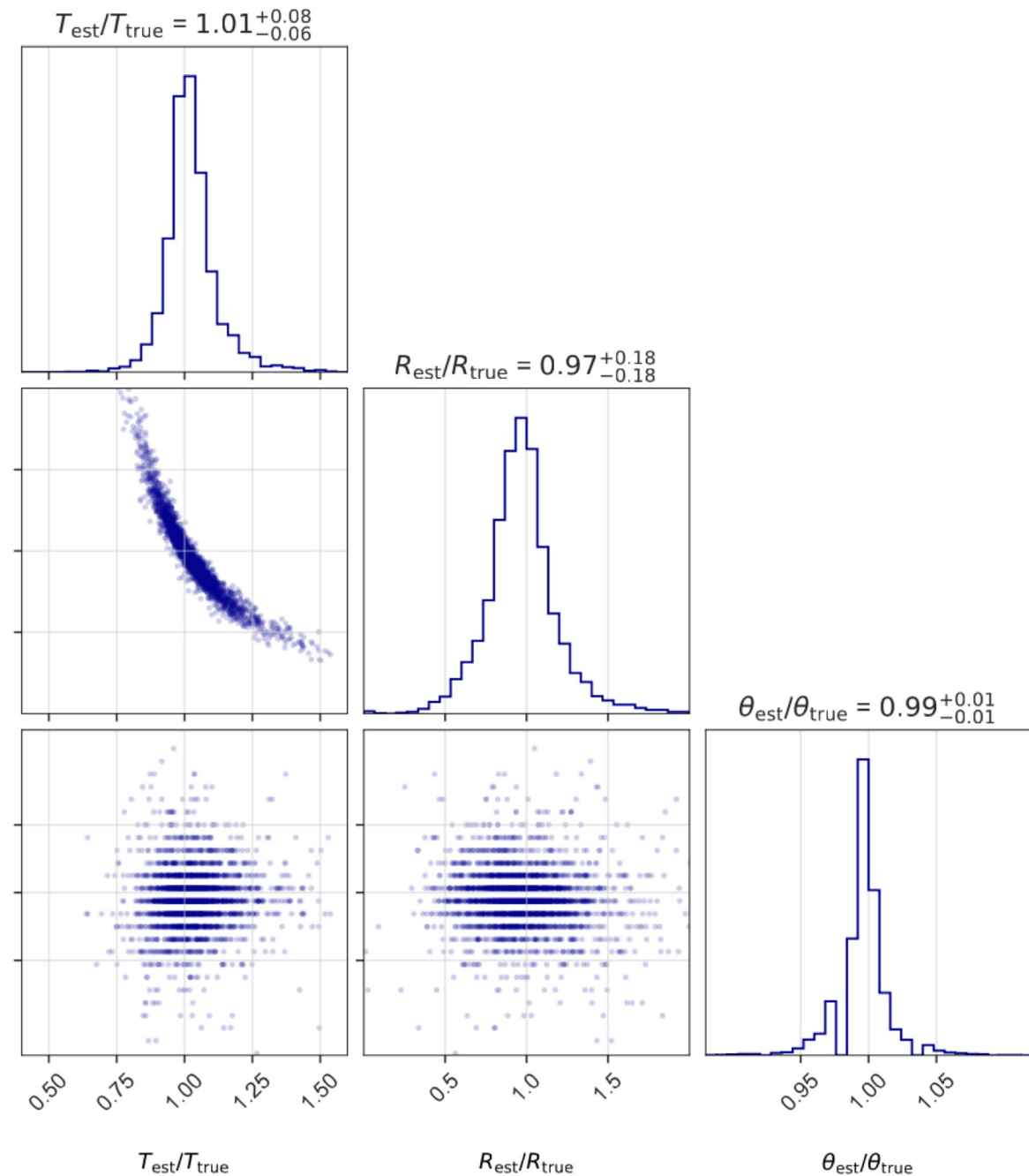


Difference map is antisymmetric wrt central point and filters out point-symmetric emission, but offset planet signal remains

Array rotation (on timescales of 16 – 20 h) will lead to a virtual path of the exoplanet emission through the difference map

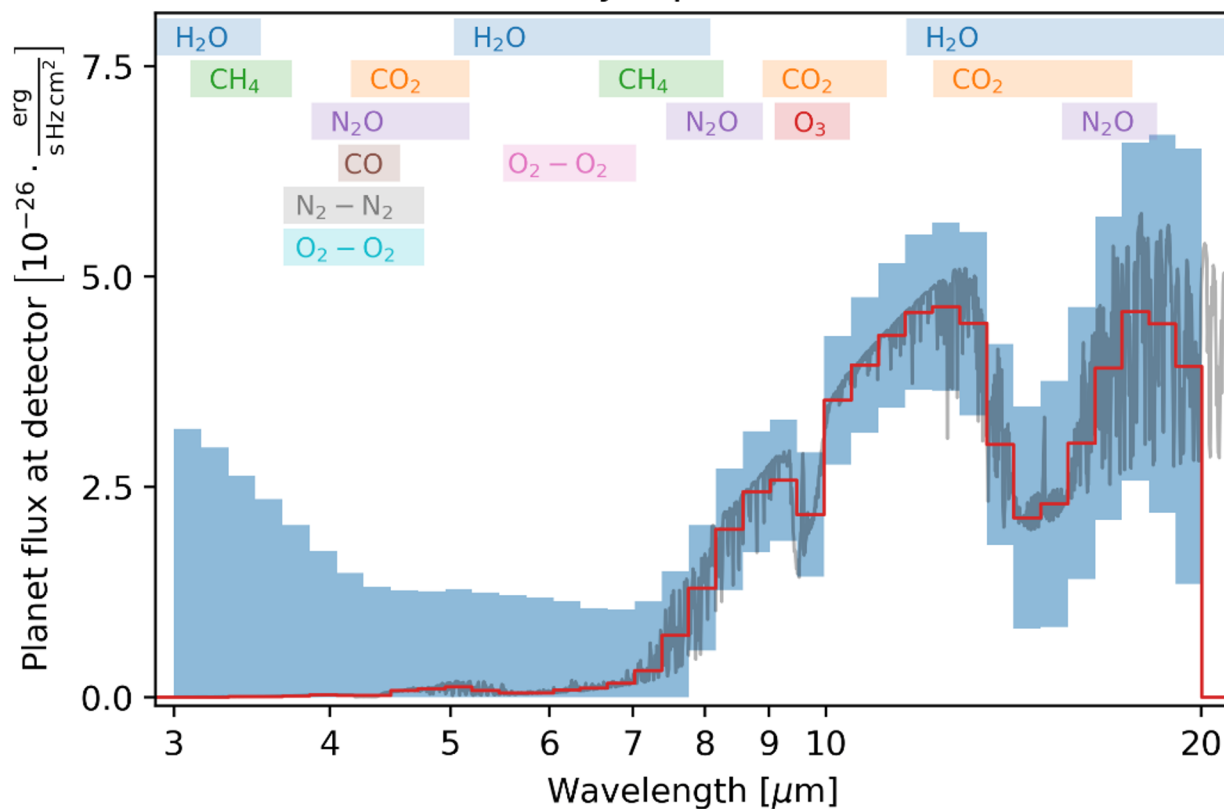
Fundamental planet parameter from single epoch

- Investigating rocky, HZ planets detected during search phase
- Signal is extracted from noisy time series and data is fitted with black-body
- Average error on
 - Temperature: ~10%
 - Radius: ~20%
 - Separation: ~1-2%

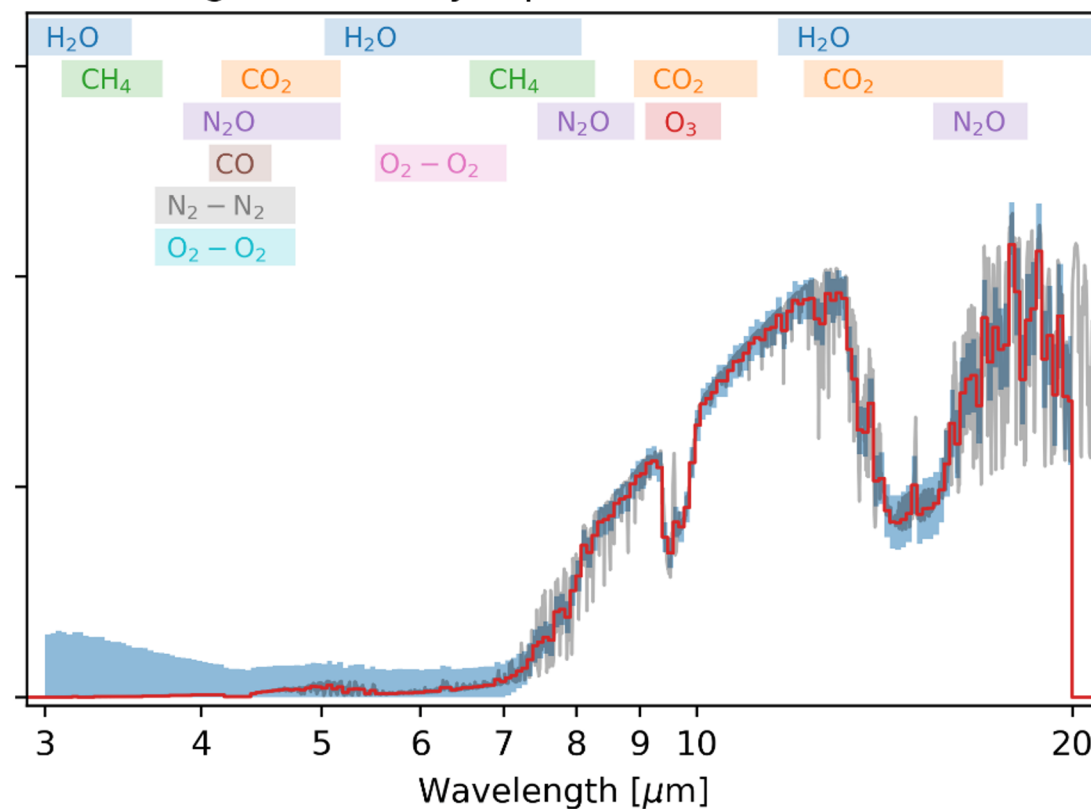


Earth-twin retrieval studies to determine characterization potential

Lowest Quality Input: $R = 20$, $S/N = 5$



Highest Quality Input: $R = 100$, $S/N = 20$

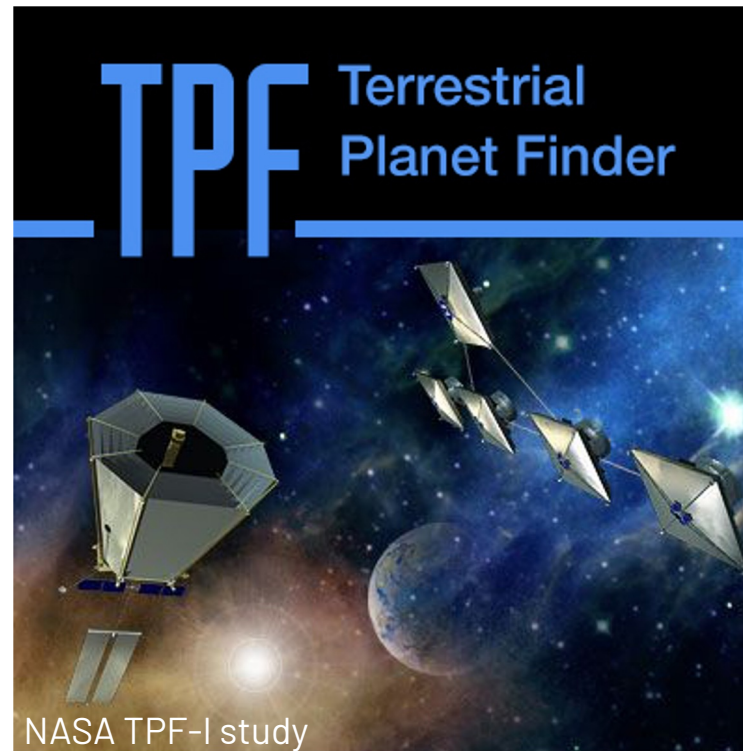


— High R Earth spectrum — Input data for retrieval ■ Uncertainty on input

Heritage

Space based (MIR, nulling) interferometry is not a new idea. However,

- Our knowledge about exoplanets has significantly increased with hundreds of terrestrial planets waiting to be discovered
- Tremendous progress was made in several key technologies



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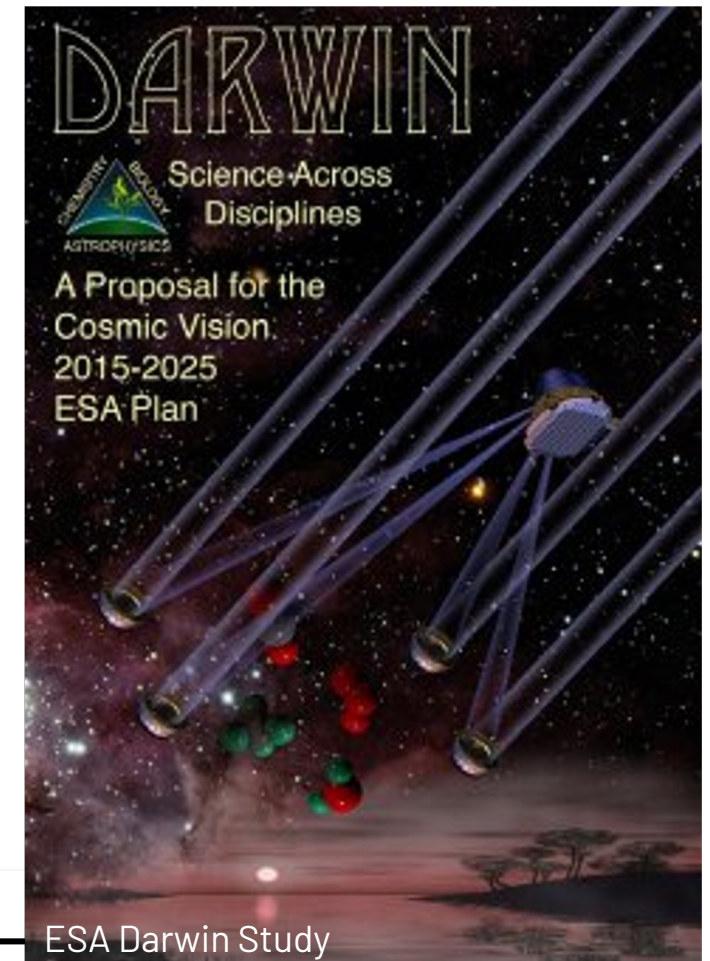
[Published: 24 August 1978](#)

Detecting nonsolar planets by spinning infrared interferometer

[R. N. BRACEWELL](#)

[Nature](#) **274**, 780–781 (1978) | [Cite this article](#)

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UV / Optical / NIR

MIR

